

## Perchlorate Formation by Electrochemistry in Martian Dust Events –II

Yan Y. C.<sup>1</sup>, Alian Wang<sup>1</sup>, J. Houghton,<sup>1</sup> Andrew Jackson,<sup>2</sup>

<sup>1</sup> Dept. Earth and Planetary Sciences and McDonnell Center for the Space Sciences, Washington University in St. Louis, 63130, USA, <sup>2</sup>Civil and Environmental Engineering, Texas Tech University, 911 Boston, Lubbock, TX 79409-1023, USA, E-mail: ycyan@levee.wustl.edu

We have reported early<sup>1,2,3</sup> that the electrochemistry processes, through Mars atmosphere to surface interaction induced by Mars Dust Storm and Dusts Devil, can be an important mechanism to form oxychlorine salts ( $\text{Cl}^{+1}$ ,  $\text{Cl}^{+3}$ ,  $\text{Cl}^{+5}$ ,  $\text{Cl}^{+7}$ ), which would be extremely important for Cl-cycle on Mars. This abstract reports further in-depth electrostatic discharge (ESD) experimental studies on Cl-phase transformations, started from three chloride salts (Na, Mg, Ca), with or without structural  $\text{H}_2\text{O}$ , with an emphasis on the yields of (per)chlorates, their depth-dependence, and the potential intermediate Cl-O-species during Mars dust events.

We built an apparatus in our Planetary Environment and Analysis Chamber (PEACH)<sup>1,4</sup>, and realized stable ElectroStatic Discharge (ESD) in the form of Normal Glow Discharge (NGD). The PEACH is capable of maintaining Mars atmospheric pressure, composition (pure  $\text{CO}_2$ ,  $\text{CO}_2+\text{H}_2\text{O}$ , and Mars Simulating Gas Mixture), and a well-controlled sample temperature range relevant to Mars surface and shallow subsurface, and with four *in situ* sensors for the characterization of molecular species as solid, liquid, and of atomic & ionic species in ESD-stimulated plasma.

We found (1) with the same electron flux density and the same ESD time duration, the perchlorate yields from a set of different ESD sample cells are proportional to the effective electrode area ratio (100:31:3.4), which suggests that the electron avalanches generated by ESD-NGD have extremely large number of electrons to react with the chloride molecules with high productivity; (2). The types of cation in starting chloride salts does have an effect on the yield of (per)chlorates,  $\text{MgCl}_2$  transfers to  $\text{Mg}(\text{ClO}_3)_2$  and  $\text{Mg}(\text{ClO}_4)_2$  more readily than the similar transformation from  $\text{NaCl}$  and  $\text{CaCl}_2$  to relevant (per)chlorates; (3) an obvious surface enrichment of  $\text{ClO}_3$  in all experiments indicating an atmosphere-surface interaction; (4) the existence of structural  $\text{H}_2\text{O}$  in the starting species influenced the yield of (per)chlorates; (5) the yield of  $\text{ClO}_4^-$  is much lower than that of  $\text{ClO}_3^-$  when started from chlorides, but much higher when the starting phase is  $\text{NaClO}_3$ , implying a sequential generation of  $\text{ClO}_4^-$  is possible on Mars. The next step of our study is ongoing in two aspects: (a) catalysis effect of co-existing species and (b) Cl-isotopic differentiation.

[1] Wu & Wang, 2016, LPSC, #2227; [2] Yan et al., 2017, LPSC, 2413; [3] Wang et al., 2017, LPSC, #2685; [4] Sobron & Wang, 2011, J. Raman Spectroscopy, DOI 10.1002/jrs.3017.

**Acknowledgments:** The authors thank NASA MoO project (06-Scout06-0027-#49137- NRA 1295053) for ExoMars mission, and NASA SSW program supported project (80NSSC17K0776).