# The Role of H<sub>2</sub>O in Subduction Zone Magmatism

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#### Keywords

volcanic arc, subducted slab, chlorite, H<sub>2</sub>O-saturated, hydrous magma, arc magmas

#### Abstract

Water is a key ingredient in the generation of magmas in subduction zones. This review focuses on the role of water in the generation of magmas in the mantle wedge, the factors that allow melting to occur, and the plate tectonic variables controlling the location of arc volcanoes worldwide. Water also influences chemical differentiation that occurs when magmas cool and crystallize in Earth's continental crust. The source of  $H_2O$  for arc magma generation is hydrous minerals that are carried into Earth by the subducting slab. These minerals dehydrate, releasing their bound  $H_2O$  into overlying hotter, shallower mantle where melting begins and continues as buoyant hydrous magmas ascend and encounter increasingly hotter surroundings. This process is controlled by plate tectonic variables that ultimately influence the location of the active volcanic arc above subduction zones. Water also modifies the thermodynamic properties of melts, leading to the unique chemical composition of arc volcanic rocks and Earth's continental crust.

#### **INTRODUCTION**

#### Pressure and temperature (P-T):

refers to the conditions at which a rock or melt last equilibrated with each other or their source region

#### Fractional

crystallization: the process of changing a liquid composition by noncongruent crystallization and immediate and permanent chemical separation of the crystalline products from the residual liquid

**Primary magma:** a melt that has not been modified chemically in any manner since it last equilibrated with its source region

For inhabitants of Earth, the occurrence of volcanic island arcs and the presence of a SiO<sub>2</sub>-rich continental crust seems natural. But these features are truly rare on other terrestrial planets, where basalts are the dominant form of volcanic material that cover the planet's surface. On Earth, SiO<sub>2</sub>-rich arc magmas that ultimately form the continental crust are a consequence of  $H_2O$  influencing melting in the subduction zone environment. This paper discusses the role of  $H_2O$  in magma-generation processes at subduction zones. The goal is to provide an updated view of melting from the bottom of the mantle wedge to the top of the crust of the overlying plate. We begin by discussing constraints on the pressure-temperature (P-T) conditions in the mantle wedge. We then discuss the way in which  $H_2O$  catalyzes melting in the mantle and the amount of  $H_2O$  involved in this melting process, before discussing the processes that occur when fractional crystallization takes place in the crust in the presence of  $H_2O$ . We conclude by discussing how plate tectonics influences the melting process to control the location of volcanoes in island and continental arc chains (for further details, see sidebar, Volcanic Arcs).

# CONSTRAINTS ON THE PRESSURE-TEMPERATURE CONDITIONS IN THE WEDGE AND SUBDUCTED CRUST

#### **Evidence from Primary Arc Magmas**

Experimentally determined temperatures of model primary arc magmas provide the best evidence of mantle temperatures. Truly primary magmas are samples of partial-melting processes in the mantle that have been delivered to the surface without any modification by fractional crystallization. Melts that have remained unaffected by modification processes are rarely found, but primitive arc magmas that are near-primary melts are present. Primitive magmas are inferred to have Mg# > 0.7 ( $Mg\# = Mg/[Mg + Fe^{2+}]$  molar) and be in equilibrium with >Fo<sub>90</sub> olivine in the mantle. Lavas with these characteristics have been identified in several arc settings (NE Japan, Mexico, Cascades, Indonesia, and Tonga) (Kelemen et al. 2003, Gaetani & Grove 2003, Hesse & Grove 2003, Wood & Turner 2009), but they are not commonplace. The chemical compositions of primitive, near-primary arc lavas and their experimentally determined pressures of last equilibration with the mantle and pre-eruptive H<sub>2</sub>O contents are found in Table 1. One important characteristic shared by all mantle-derived melts found in arcs is that their last equilibration in the mantle appears to have occurred at  $\sim 1$  GPa and 1100–1500°C, which represents conditions near the crust-mantle boundary in continental arcs. This shallow depth of equilibration provides important evidence to constrain models of mantle melting in arcs. We return to these matters in a later discussion of variations in the chemical composition and water contents of arc magmas. Another important characteristic of primitive arc magmas is their range in chemical composition from low-SiO<sub>2</sub> high-alumina olivine tholeiite to basaltic andesite and primitive magnesian andesite (Table 1), which is compositionally distinct from the more abundant andesite and dacite

#### **VOLCANIC ARCS**

Often abbreviated as arc, volcanic arc refers to a chain of volcanoes often forming an arcuate shape in map view. Such chains form at convergent tectonic-plate boundaries above a subducting oceanic plate. The volcanoes form approximately parallel to the oceanic trench, where the two plates meet, and is considered an "oceanic" or "island" arc when the overriding plate is oceanic and a "continental" arc when the overriding plate is continental.

lavas
near-primary arc
primitive,
compositions of
Chemical
Table 1

															FeO/	
Location <sup>a</sup>	$^{\rm qL}$	$\mathbf{P}^{\mathrm{b}}$	$\mathrm{H}_2\mathrm{O}^\mathrm{b}$	$SiO_2$	$TiO_2$	$M_2O_3$	FeO	MnO	MgO	CaO	Na <sub>2</sub> O	$\mathbf{K}_2\mathbf{O}$	$P_2O_5$	Mg#	MgO	<b>Reference</b> <sup>c</sup>
TGIJ	1125	1.2	7.5	59.59	0.44	13.55	6.32	0.12	9.65	0.24	2.66	1.3	0.13	0.73	0.65	1
SD-438 J	1300	1.2	0	49.76	1.02	15.48	8.87	0.16	11.68	8.91	2.6	1.28	0.22	0.7	0.76	2
SH-85-41c C	1180	1	9	57.79	0.6	14.46	5.74	0.11	9.14	8.17	3.11	0.71	0.15	0.74	0.63	3
SH-85-44 C	1200		4.5	51.68	0.6	16.4	7.93	0.16	10.79	9.67	2.24	0.42	0.11	0.71	0.73	4
ML-82-72f C	1310	1.1	0	47.4	0.59	18.5	8.2	0.15	10.52	12.02	2.16	0.07	0.06	0.7	0.78	5
M.102 M	1300	1.7	6	49.4	0.99	14.84	7.63	0.12	11.63	8.3	3.53	1.4	0.44	0.73	0.66	6
Av-MGB I			0.3	49.1	0.82	15.9	8.9	0.17	10.5	11.3	2.21	0.34	0.1	0.68	0.85	7
BON IB	1480	1.5	2.5	51.88	0.7	8.71	8.94	0.12	20.69	7.1	1.7	0.16		0.8	0.43	8
V35-5 A				56.54	0.61	16.68	4.84	0.12	6.62	7.2	3.89	1.3	0.13	0.75	0.73	6
ID16A	1320	1.2	0	48.9	0.7	16.01	8.9	0.17	11.42	10.89	2.21	0.52	0.12	0.7	0.78	10

<sup>a</sup>Key to locations: J, Japan; C, Cascades; M, Mexico; I, Indonesia; IB, Izu-Bonin; A, Aleutians.

<sup>b</sup>T (temperature) in °C; P (pressure) in GPa;  $H_2O$  contents in weight percent (wt%).

<sup>c</sup>References: 1. Tatsumi (1981). 2. Tatsumi et al. (1983). 3, 4. Grove et al. (2003). 5. Bartels et al. (1991). 6. Hesse & Grove (2003). 7. Sisson & Bronto (1998). 8. Falloon & Danyushevsky (2000). 9. Yogodzinski et al. (1995). 10. Draper & Johnston (1992). lavas that make up the bulk of island and continental arc volcanoes. We also return to a discussion of  $H_2O$ 's influence on these characteristics later in the paper.

#### **Metarodingite:**

a metamorphosed and/or metasomatized rodingite, which is an altered oceanic basalt that tends to be rich in grossular garnet and calcic pyroxene enveloped in serpentinite

**Slab:** a subducting oceanic plate composed of sediment, basaltic oceanic crust, and lithospheric mantle as well as the aqueous alteration of these constituents

#### **Evidence from Subducted Oceanic Crust**

P-T conditions preserved in exhumed subducted oceanic lithosphere would provide the ideal evidence for the conditions where the transfer of chemical components in either fluid and/or melt takes place in subduction zones (pressure of 3 GPa or greater). Only a handful of localities are known to preserve rocks that experienced these conditions: the Western Gneiss region, Norway (Hacker et al. 2010); the Sanbagawa belt, SW Japan (Mori & Banno, 1973); and Cima di Gagnone in the Swiss Alps (Evans & Trommsdorff 1974, 1983). The Sanbagawa subduction-type metamorphic belt in SW Japan represents the deepest exposed portion of a Mesozoic accretionary complex along the Japanese island arc. Located on the island of Shikoku, the Higashi-akaishi peridotite body is the largest ultramafic lens within the Sanbagawa belt and is dominantly composed of dunite, lherzolite, and garnet clinopyroxenite, interfingered in one locality with quartz eclogite (Enami et al. 2004, Mizukami & Wallis, 2005). Previous work indicates the P-T history of the peridotite includes prograde metamorphism with peak temperatures of 700-810°C and pressures of 2.9-3.8 GPa at approximately 110-120 million years ago (Enami et al. 2004). In the Swiss Alps, rocks exposed in the Cima di Gagnone, Finero (Zanetti et al. 1999, Brouwer et al. 2005), and Alpe Arami (Brenker & Brey 1997, Paquin & Altherr 2001, Olker et al. 2003, Nimis & Trommsdorff 2001, Bocchio et al. 2004) regions hold the potential to reveal mechanisms of hydrous mantle melting. Previous work on the Gagnone metamorphic body (Evans & Trommsdorff 1974, 1983; Pfiffner & Trommsdorff 1998) indicates it experienced peak metamorphic conditions of 2.4 GPa and 800°C and contains lenses of garnet-bearing peridotite, chlorite peridotite, and eclogite folded between pelitic gneisses. Metarodingites are found in association with the eclogite and suggest at least a portion of this ultramafic body preserves a remnant of oceanic crust. The peak-metamorphic conditions of these ocean crust and mantle-wedge-related metamorphic rocks place them at a depth near the wedge-slab interface-the subducted lithosphere boundary that is below the location of volcanic arcs.

#### Models of the Thermal Structure of Subduction Zones

Beginning with the classic study of Toksoz et al. (1971), numerical modelers have sought to infer the physical conditions encountered when subduction of the cold oceanic lithosphere causes the flow of mantle material into and around the wedge corner and back down into the mantle. Early models reproduced the basic thermal structure, but could not recover exact conditions inferred from petrologic constraints. Models have greatly improved in the past 10 years (van Keken et al. 2002, Kelemen et al. 2003, Cagnioncle et al. 2007) and now predict mantle wedge conditions that approach those inferred from the petrologic evidence (Figure 1). Notable progress has been achieved by understanding the depth at which coupling occurs between the subducted slab and the mantle wedge (e.g., Wada & Wang 2009) and by implementing more sophisticated models of temperature-dependent rheology in mantle flow (e.g., van Keken et al. 2008). Syracuse et al. (2010) developed a comprehensive model that spans the global variability in subduction zone physical parameters and predicts the thermal structure of the wedge-lithosphere systems. The temperatures predicted at the base of the mantle wedge (specifically the mantle directly above the interface between the wedge and the subducted oceanic lithosphere) from these models are also shown in Figure 1. Here we use the results of these new models of P-T conditions in the mantle wedge and subducted lithosphere to discuss the petrologic controls on melting involving water.





Pressure-temperature diagram comparing the conditions of the last equilibration with the mantle of near-primary arc magmas (green dots; see Table 1) and the range of temperatures and pressures predicted by models of the physical processes occurring in the mantle wedge in arc environments (gray boxes). The field labeled "Hottest shallow nose of wedge" is the part of the wedge where the highest hydrous flux melting temperatures are reached (e.g., to the backarc side of path B in Figure 6a labeled HNW). "Conditions at base of mantle wedge" refers to the temperature at the contact between the subducted oceanic lithosphere and the overlying mantle wedge (e.g., base of dashed line in **Figure 6***a*) where fluid-fluxed melting is occurring.

#### MELTING PROCESSES IN THE ARC MANTLE

It is generally agreed that melting in subduction zones is initiated in the mantle wedge above the subducted oceanic lithosphere due to an influx of volatiles (Gill 1981, Tatsumi et al. 1986, Davies & Stevenson 1992, Gaetani & Grove 2003). We propose that this melting occurs at low temperatures when a small amount of  $H_2O$ -rich fluid causes the first melt to be generated at the vapor-saturated solidus (P1,T1 in Figure 2). However, no consensus exists regarding the specific nature and location of this melting process.

In contrast, mantle melting beneath mid-ocean ridges occurs by adiabatic decompression under anhydrous conditions (Klein & Langmuir 1987, Asimow et al. 2001) (Figure 2) (anhydrous adiabat). Because the adiabat has a shallower slope than does the dry peridotite solidus, upwelling mantle intersects and crosses the solidus and obtains the heat of fusion by cooling from the adiabat to the melting curve. This process has been proposed to operate in the subduction environment (Tatsumi et al. 1983, Nye & Reid 1986, DeBari et al. 1987, Plank & Langmuir 1988), and there are examples of near-anhydrous melts that have been generated in subduction zones (Tatsumi et al. 1983, Bartels et al. 1991, Sisson & Bronto 1998, Elkins-Tanton et al. 2001) by mantle that is transported into the mantle wedge by corner flow (Wiens et al. 2008).

Mechanisms other than adiabatic decompression melting are required to form hydrous magmas. For example, in hydrous-flux melting, incipient melting begins near the base of the mantle H<sub>2</sub>O- or vapor-

saturated: describes a liquid that is in equilibrium with water in excess of that which can be dissolved in the liquid

Anhydrous: lacking  $H_2O$ 



This pressure-temperature diagram shows three styles of mantle-melting processes. The anhydrous (*red line*) and  $H_2O$ -saturated (*dark blue line*) melting solidi for peridotite are shown. In adiabatic decompression melting under anhydrous conditions (*orange arrow*), the melt and mantle cool as they ascend to shallower depths. Melting by this mechanism, if it involved a vapor-saturated first melt on the  $H_2O$ -saturated solidus, would crystallize the melt and place it in the vapor + crystals field. Hydrous flux melting (*light blue arrow*) can occur in the mantle wedge because the temperature increases above the slab-wedge interface. A vapor-saturated melt (*light blue arrow*) produced by this process ascends into shallower hotter mantle and comes into thermal and chemical equilibrium by increasing temperature and melt fraction. After Gaetani & Grove (2003).

wedge in the presence of a small amount of  $H_2O$  derived from the slab or wedge, because the presence of  $H_2O$  depresses the peridotite solidus. If this melt and its matrix were to ascend into the overlying wedge along an adiabatic gradient, it would enter the crystal plus vapor field and freeze (**Figure 2**, hydrous adiabat) within the mantle wedge (Nicholls & Ringwood 1972, Tatsumi & Eggins 1995). When excess  $H_2O$  is present, melting can continue only if the ascending mass is small enough that it can be thermally equilibrated by the surrounding hotter mantle and heat can be transferred into it to drive a melting process. Potential melting mechanisms that would allow this to happen include the following: (*a*) the development of small diapirs of less dense, partly molten mantle that rise into the mantle wedge and further melt as they decompress (Tatsumi et al. 1983, 1986; Kushiro 1990; Pearce et al. 1995) or (*b*) a reactive porous-flow process (Kelemen 1986; Kelemen et al. 1990; Grove et al. 2002, 2006). In reactive porous flow (labeled hydrous flux melting in **Figure 2**), a buoyant H<sub>2</sub>O-rich melt percolates up through the mantle wedge and reacts with and dissolves the surrounding mantle, thus decreasing the relative abundance of H<sub>2</sub>O in the melt.

When solid mantle and a small amount of  $H_2O$ -rich fluid are heated to temperatures in excess of the vapor-saturated solidus, melt is generated (P<sub>1</sub>, T<sub>1</sub>, **Figure 3**). This first melt is extremely rich in  $H_2O$  (~30 wt%) (Hodges 1974, Mysen & Boettcher 1975) and will be positively buoyant



A polybaric temperature-composition phase diagram that illustrates the process of hydrous flux melting of the mantle wedge by reactive porous flow or by ascent of small, but buoyant, low-viscosity diapirs of mantle plus melt. The temperature-pressure path followed during this melting process is shown in **Figure 2** as "Hydrous flux melting" and the melting relations for two pressures are shown. At P<sub>1</sub>, T<sub>1</sub>, melting begins at the vapor-saturated solidus. This fluid-saturated melt ascends to shallower depths (P<sub>2</sub>, T<sub>2</sub>) in the inverted temperature gradient of the mantle wedge, where it is out of equilibrium with its surrounding mantle. The melt contains more H<sub>2</sub>O than does an equilibrium melt at the same temperature, and the melt dissolves the surrounding mantle, increasing melt fraction and decreasing melt H<sub>2</sub>O content. Modified after Grove et al. (2002).

and percolate upward by porous flow. At these shallower, hotter conditions, the melt will be vapor undersaturated and superheated with respect to the vapor-saturated solidus, placing it out of equilibrium with its surroundings. The melt will thus tend to dissolve the surrounding silicate minerals in the peridotite and approach chemical and thermal equilibrium causing the H<sub>2</sub>O content of the melt to decrease ( $\Delta$ F) (**Figure 3**) until it reaches the equilibrium concentration of the vapor-undersaturated liquid-solid boundary at (P<sub>2</sub>, T<sub>2</sub>). This process will continue as the melt ascends to shallower depths and encounters progressively hotter peridotite. In this porous-flow model, the melt composition and H<sub>2</sub>O content are dictated by the last depth of equilibration.

Models of diapiric flow have also been proposed (Hall & Kincaid 2001). In these models, slab rollback leads to three-dimensional mantle-flow patterns and the development of diapirs that ascend into the hotter mantle and can be melted. Melts generated through this mechanism are likely to be batch melts that re-equilibrate with the diapir mass as they ascend. By contrast, in reactive porous-flow, the ascent velocity of the melt and the permeability of the mantle will exert controls on the time that melt and mantle remain in contact and thus effect the composition of melts formed (Grove et al. 2002).

#### MANTLE PHASE EQUILIBRIA

#### Peridotite-H<sub>2</sub>O Melting

Between 1968 and 1975, four experimental studies reported on the melting of peridotite under  $H_2O$ -saturated and upper-mantle pressure conditions (Kushiro et al. 1968, Green 1973, Millhollen et al. 1974, Mysen & Boettcher 1975). These studies were revisited by Kawamoto & Holloway (1997) and more recently by Grove et al. (2006), Green et al. (2010), and Till et al. (2011), and a thermodynamic model of  $H_2O$ -saturated melting was presented by Smith & Asimow (2005). The P-T conditions of  $H_2O$ -saturated melting in the early work fell into one of two groups depending on the duration of the experiment. Shorter experimental durations (<3 h) found solidus temperatures of ~1000°C at 3 GPa while longer experimental durations (24–64 h) yielded solidus temperatures of a fertile peridotite composition and reported a low-temperature solidus (**Figure 4**) (800°C at 3 GPa) similar to that obtained by Mysen & Boettcher (1975). Green et al. (2010) reported that melting begins at a higher temperature (1100°C at 3 GPa) (**Figure 5**) similar to their earlier work (Green 1973). The difference between the two studies lies in the interpretation of what constitutes the first melt. As the diagnostic criterion for the presence of quenched hydrous



#### Figure 4

Experimentally determined phase diagram for  $H_2O$ -saturated peridotite from Grove et al. (2006) and Till et al. (2011). Experiments were conducted with the primitive mantle composition of Hart & Zindler (1986). Hydrous and aluminous phase stabilities that are not constrained by the experiments conducted by Grove et al. (2006) and Till et al. (2011) are derived from Ulmer & Trommsdorff (1995), Fumagalli & Poli (2005), and Pawley (2003). Abbreviations: amph, amphibole; atg, antigorite; Chl, chlorite; Cpx, clinopyroxene; tlc, talc.

#### **QUENCH MATERIAL**

At the termination of an experiment, the sample is cooled rapidly by turning off the furnace power. The hydrous silicate melt and/or fluid portions of the experiment do not remain a single homogenous material, but instead decompose to a mixture of rapidly grown crystals and glass referred to as "quench material."

silicate melt in an experiment, Green (1976) and Green et al. (2010) use the presence of interstitial patches of amphibole and/or quench clinopyroxene, mica, and glass all with Mg# of 70–85 within the experiment. Conversely, the lack of mafic silicates in the quench material is interpreted as a quench of a subsolidus vapor phase (see sidebar, Quench Material). Till et al. (2011) show that above 800°C at 3–4 GPa, and until 1100°C at 3 GPa, the Mg# of crystalline phases changes in a manner that is consistent with solid-melt partitioning rather than solely solid-fluid partitioning. Thus, the differences in the two interpretations might be a matter of semantics, and can be resolved once the "melt" or "fluid" compositions are directly analyzed.

#### Hydrous Mineral Phase Stability

Chlorite is stable to temperatures just above the H<sub>2</sub>O-saturated solidus between 2 and 3.6 GPa (Pawley 2003, Fumagalli & Poli 2005, Grove et al. 2006, Till et al. 2011, Dvir et al. 2011) (**Figure 4**). Above 3.6 GPa, the H<sub>2</sub>O-saturated solidus and the chlorite stability fields diverge. Chlorite contains ~13 wt% H<sub>2</sub>O in its structure, and a fertile peridotite has enough Al<sub>2</sub>O<sub>3</sub> to stabilize 6–7 wt% chlorite. This translates to ~2 wt% bulk H<sub>2</sub>O in meta-peridotite at the P-T conditions of the vapor-saturated solidus. Schmidt & Poli (1998) and Grove et al. (2006) noted that



#### Figure 5

Experimental and theoretical determinations of the H2O-saturated peridotite solidus.

#### Hydrous phase:

a mineral or melt that contains hydrogen as an essential structural constituent. In the mineral it is usually incorporated as hydroxyl (OH<sup>-</sup>) groups chlorite is a stable mineral in the P-T range appropriate for the base of the upper mantle wedge. Chlorite, as well as serpentine, likely form in the shallow part of the slab-wedge interface from fluids released during the dehydration of serpentine and amphibole in the subducting oceanic crust and lithosphere (Schmidt & Poli 1998, Iwamori 1998, Pawley 2003, Grove et al. 2006, Iwamori 2007) or from  $H_20$  produced by the dehydration of subducted sediments. Seismic images of the Honshu subduction zone reveal a low-velocity layer that is interpreted to reflect the presence of hydrous phases in the shallow wedge (Kawakatsu & Watada 2007). Thus hydrous phases, such as chlorite and serpentine, can be exposed to downward transport in the mantle wedge by corner flow.

Chlorite in fertile peridotite will breakdown at temperatures above the H<sub>2</sub>O-saturated solidus at 800 to 820°C over the pressure range of 2 to 3.6 GPa (**Figure 4**). Structural water released from chlorite can trigger melting of the mantle and/or perpetuate H<sub>2</sub>O-saturated melting previously initiated by H<sub>2</sub>O fluxing following the breakdown of other hydrous phases in the subducting mantle lithosphere. We suggest that H<sub>2</sub>O is fluxed into the base of the mantle wedge from the dehydration of phases in both the underlying subducted slab and the base of the mantle wedge (**Figure 6***a*). The existence of chlorite in the mantle wedge at temperatures 100–150°C higher than serpentine dehydration assures that H<sub>2</sub>O is available at the H<sub>2</sub>O-saturated solidus at subarc



#### Figure 6

Two diagrams that schematically illustrate the conditions that limit hydrous flux melting in subduction zones. (*a*) A typical cross section through a subduction zone using the thermal model from Grove et al. (2009) for a slab dip of  $30^{\circ}$  and a convergence rate of 40 km per million years. Also shown are the stability limits for high-H<sub>2</sub>O (>10 wt%) minerals. The melting model from Grove et al. (2006) is used to predict the extents of hydrous flux melting in the bottom half of the wedge. HNW denotes the hottest, shallowest nose of the mantle wedge where the maximum amount of melting occurs. The part of the mantle wedge directly above the interface between subducted oceanic lithosphere and the mantle is denoted SWI and extends from path A to path C. (*b*) Phase diagram from **Figure 4** with the temperature-depths paths from panel *a* superimposed. The dark gray region labeled P,T<sub>melt</sub> shows the region of pressure-temperature space where the stability of hydrous phases in the base of the mantle wedge allows H<sub>2</sub>O-saturated melting to begin. Figure with minor modifications from Grove et al. (2009). Abbreviations: HNW, hot nose of the wedge; OC, oceanic crust; SWI, slab-wedge interface.

depths. Therefore, the presence of chlorite in the wedge provides a mechanism for transporting significant quantities of water (up to 2 wt%  $H_2O$  in bulk peridotite) into the mantle, even in hot subduction zones, and removes the need for anhydrous slab melting in hot-slab environments, where the slab is thought to have lost water and must heat up to a much higher temperature before it can melt. Sources of  $H_2O$  will be present in minerals above and below the subducted oceanic crust.

The maximum depth stability of hydrous minerals, and thus the maximum depth at which water released within the slab is released into the overlying mantle wedge, is influenced by the temperature structure in the subducted lithosphere (**Figure 6***a*). In addition to chlorite, serpentine and the 10-Å phase are two hydrous minerals that are stable to considerable depths in the subducted oceanic lithosphere (up to 150 km, depending on the P-T structure of the slab) (Iwamori 1998, Pawley 2003, Fumagalli & Poli 2005). Amphibole was once proposed as a carrier for transporting water to depth in the mantle (Wyllie 1982, Davies & Stevenson 1992), but the depth to the slab beneath arc volcanoes is beyond the high-pressure stability of amphibole. As shown in recent vapor-saturated peridotite melting experiments (Grove et al. 2006, Médard & Grove 2006, Till et al. 2011), amphibole is stable from ~0.8 to 2 GPa on the solidus, and chlorite is stable from 2 to 3.6 GPa (**Figure 6***b*), although uncertainty remains about whether the experimental bulk-water content affects amphibole stability (Green et al. 2010).

Once low-extent H<sub>2</sub>O-rich melts are formed in the mantle wedge near the slab-wedge interface and ascend buoyantly into the mantle wedge, the next important petrologic control on mantle melting at subduction zones is the temperature structure in the overlying mantle wedge. In Figure 6, melting processes in the overlying wedge are illustrated by selecting three P-T paths through the region where the mantle will be saturated with an H2O-rich fluid and/or it will be hot enough to begin melting at the vapor-saturated peridotite solidus. In path A, the hottest part of the mantle wedge is at the vapor-saturated solidus, and melting can occur because of the breakdown of several hydrous minerals in the subducted lithosphere. However, as soon as that melt separates and ascends into the overlying cooler mantle, it will freeze because the P-T path drops below the vapor-saturated solidus. For path B, an H<sub>2</sub>O-rich fluid is supplied by the breakdown of chlorite from the underlying subducted lithosphere, and melting begins at the vapor-saturated solidus. We propose that the amount of melt produced will be limited by the amount of H<sub>2</sub>O added and that, generally, the supply of H<sub>2</sub>O will be a limiting factor in melting. This initially vapor-saturated melt will ascend into the overlying, hotter, shallower mantle wedge and undergo flux melting as it equilibrates with ambient temperature conditions. It will then follow a vapor-undersaturated melting path through P-T space (described above and illustrated in Figure 2). As the melt ascends into shallower mantle, it reacts and dissolves solid minerals, increasing the amount of melt and lowering the H<sub>2</sub>O content of the melt (Figure 6a shows a melting model from Grove et al. 2006). The amount of melting in a vertical column depends on the maximum temperature achieved in the hottest part of the mantle wedge and on thermal equilibration between melt and surrounding mantle, and the extent of melting will increase for vertical paths that are farther from the wedge corner. In path B, the maximum temperature and the maximum amount of melting will occur in the hot core of the wedge, and the melt will then ascend into the cooler, shallower mantle above and ultimately into the crust.

For ascent paths further away from the wedge corner (path C, **Figure 6**),  $H_2O$ —the essential ingredient to producing vapor-saturated melting—will ultimately be exhausted as the subducted lithosphere and lower part of the mantle wedge warm to temperatures above the maximum stability temperature and pressure of hydrous minerals (there are other hydrous minerals that remain stable, but their abundance is low). The absence of  $H_2O$ -bearing minerals terminates vapor-saturated melting (path C, **Figure 6**). Thus, melting is possible over a region above the slab between path

A and path C that is defined by the temperature structure in the wedge and the supply of  $H_2O$  from the lithosphere and lower part of the mantle wedge.

The above discussion is a simplified explanation that focuses on the P-T constraints from phase equilibria. There are other important variables that need to be considered, including the permeability of the mantle and available pathways for fluid and melt ascent in the mantle wedge. Mibe et al. (1999) demonstrated that the permeability of an olivine-rich peridotite varies with pressure and temperature in the mantle wedge and that fluid can ascend into the wedge and allow melting to begin only when the grain-boundary geometry changes to produce an interconnected, permeable network. Cagnioncle et al. (2007) proposed that mantle flow and temperature structure in the overlying wedge can deflect the ascent paths of melt as it ascends into the overlying wedge, and Hall & Kincaid (2001) suggested that diapiric flow and the establishment of melt channels may influence the ascent of melt in the wedge. We return to these models in the last section of this review where we determine the extent to which the variability in the location of arc volcanoes can be explained by the supply of water, temperature structure, and kinematic variables associated with subduction.

#### COMPOSITIONS OF MAGMAS IN SUBDUCTION ZONES

In the Introduction, we highlight the uniqueness of Earth's SiO<sub>2</sub>-rich continental crust, which, on average, is thought to have the composition of andesite (Rudnick & Fountain, 1995). The presence of magmatic  $H_2O$  appears responsible for this SiO<sub>2</sub> enrichment in two ways: (a) The processes of hydrous mantle melting can lead to the direct production of andesitic primary magmas, and (b) fractional crystallization of H<sub>2</sub>O-bearing melts influences the mineral phase appearance sequence in basaltic and andesitic primary magmas leading to the production of SiO2-rich continental crust. Both processes produce mafic residual products from intracrustal fractional crystallization that must be returned into the mantle by some process. A mechanism that can return these mafic cumulates to the mantle is lower crustal delamination (Kay & Kay 1993, Rudnick 1995), whereby mafic cumulates sink into the lower crust and return to the convecting mantle. We suggest that these water-controlled mechanisms are the dominant ones that produce SiO2-rich continental crust, but there are other ways to make andesite and dacite melts. For example, assimilation of SiO<sub>2</sub>-rich material by hot anhydrous basaltic magma during crystallization in the crust (DePaolo 1981, Grove & Kinzler 1986) and melting of hydrated basaltic crustal materials (e.g., amphibolite) (Rushmer 1991, Sen & Dunn 1994, Rapp & Watson 1995) can also produce SiO<sub>2</sub> enrichment in magmas.

#### MANTLE-MELTING CONTROLS

An examination of **Table 1** documents that near-primary arc magmas are sometimes basaltic in composition and may be either dry or H<sub>2</sub>O rich. Melting experiments on mantle peridotite at H<sub>2</sub>O-undersaturated conditions leads to the production of melts that fall within the basalt compositional range (Kushiro 1990, Hirose & Kawamoto 1995, Hirose 1997, Gaetani & Grove 1998, Parman & Grove 2004). These mantle melts resemble primary anhydrous melts of mantle lherzo-lite, but they differ in several important characteristics. These experimentally produced melts are characterized by higher SiO<sub>2</sub>/(MgO+FeO), and the presence of 3 wt% H<sub>2</sub>O in the melt increases SiO<sub>2</sub> in the melt by 1 wt% (melt composition recalculated without H<sub>2</sub>O) (Gaetani & Grove 1998). Additionally, evidence for the involvement of H<sub>2</sub>O in arc magma melt generation is manifested in the mineral-melt partition coefficients (e.g.,  $D_{MgO}^{opx/liq} = \% MgO_{opx}/\% MgO_{liq})$  for MgO between orthopyroxene and liquid as well as between olivine and liquid. The partition coefficients of



Mineral-melt partition coefficients for orthopyroxene ( $D^{opx}$ ) in anhydrous (*solid symbols*) and H<sub>2</sub>O-bearing (*open symbols*) melts of mantle peridotite. The H<sub>2</sub>O-bearing melt compositions are recalculated on an anhydrous basis. The partition coefficient for MgO increases as temperature decreases, but the D for SiO<sub>2</sub> remains constant. The consequence is that lower temperature melts have higher SiO<sub>2</sub>. After Parman & Grove (2004).

both minerals are strongly temperature dependent, with D increasing dramatically as T decreases (**Figure 7**) and the presence of  $H_2O$  broadening the temperature range of silicate liquid stability to much lower temperatures. However, the D for SiO<sub>2</sub> remains constant, showing only a small temperature dependency. Therefore, mantle lherzolite melts become enriched in SiO<sub>2</sub> because the presence of high amounts of dissolved  $H_2O$  lowers the melting temperatures of the mantle.

This effect can be observed in the compositions of natural arc magmas compiled in **Table 1** and in the covariations in temperature,  $H_2O$  content, and  $SiO_2$  content (**Figure 8**). Both anhydrous and hydrous basaltic liquids are in equilibrium with their mantle source at high temperatures, but higher  $SiO_2$  primitive arc magmas are higher in  $H_2O$ . Melting extent and mantle temperature also play a role, and the boninite compositions included in this data set reflect the effects of hightemperature high-extent hydrous melting. The results of peridotite melting experiments at 0.75 to 2 GPa in the presence of excess  $H_2O$  exemplify these effects (Kushiro 1972, Grove et al. 2006, Till et al. 2011) and contain near solidus melts with >12 wt%  $H_2O$  at 57 to 60 wt% SiO<sub>2</sub>.

#### EFFECTS OF FRACTIONAL CRYSTALLIZATION

The FeO\*/MgO versus SiO<sub>2</sub> diagram was introduced by Miyashiro (1974), who used it to discriminate arc (calc-alkaline) from tholeiitic differentiation trends and to distinguish among different types of calc-alkaline trends that he observed in arcs in the Pacific Ring of Fire. On the Miyashiro (1974) diagram, an anhydrous fractional crystallization trend (tholeiitic trend) is nearly vertical, with FeO\*/MgO increasing continuously at constant SiO<sub>2</sub>. The differentiation trend followed during fractional crystallization of Galapagos mid-ocean ridge lavas (Grove & Kinzler 1986, **Boninite:** An arcrelated volcanic rock containing both high SiO<sub>2</sub> (>53 wt.%), high MgO (>8 wt.%) and low TiO<sub>2</sub> (<0.5 wt.%) typically formed during the early stages of subduction in the fore-arc environment



Variations in temperature of mantle melting and in the  $SiO_2$  and  $H_2O$  compositions of the near-primary arc mantle melts from **Table 1**. Abbreviation: wt%, weight percent.

Juster et al. 1989) provides an excellent example of a dry tholeiitic trend (Figure 9b). Sisson & Grove (1993a) provided the first experimental evidence that liquids produced by crystallization of hydrous basaltic magmas define crystallization paths similar to those found in many calc-alkaline suites, and these lead to high-SiO<sub>2</sub> and esitic residual liquids. The trends illustrated in Figure 9 are generated in  $H_2O$ -saturated experiments carried out at 200 MPa with ~6 wt%  $H_2O$  in the melt (equivalent to a depth of 6 km in the crust and typical conditions for fractional crystallization of calc-alkaline magmas). Under these conditions,  $H_2O$  has a significant effect on the mineral phase stability and crystallization sequence and, therefore, on the compositional evolution of liquids produced during fractional crystallization. The major effect of  $H_2O$  is to suppress the crystallization of plagioclase and put Fe-Mg-rich silicates on the liquidus. The presence of  $H_2O$  also lowers the temperature at which the silicate minerals crystallize, but it has little effect on the temperature at which spinel appears. For example, an iron-rich spinel appears at or near the liquidus of a hydrous basaltic melt (Sisson & Grove 1993a). When late-appearing plagioclase crystallizes, it is rich in calcium and aluminum, which further promotes enrichment of SiO<sub>2</sub> and alkalis in the residual liquids. Therefore, liquids produced by fractional crystallization become depleted in FeO and enriched in  $SiO_2$  to produce the calc-alkaline differentiation trend.



Variation diagrams showing FeO\*/MgO versus SiO<sub>2</sub> [all in weight percent (wt%)] and the calc-alkaline (CA) and tholeiitic (TH) dividing line of Miyashiro (1974). FeO\* implies that all of the iron in the melt is calculated as divalent iron. (*a*) The compositional range of the global average of arc lava suites compiled by Ewart (1979, 1982) and suites of low FeO\*/MgO from Mt. Shasta, California; the Miocene Setouchi volcanic belt, Japan; and Adak Island in the Aleutians (*red squares*). (*b*) The primary mantle melts from **Table 1** (*gray stars*) define the mantle melting array for global arc suites. This trend extends from low-SiO<sub>2</sub> melts in equilibrium with a lherzolite mantle residue (olivine + orthopyroxene + high-Ca pyroxene + spinel) to high-SiO<sub>2</sub>, high-degree melts in equilibrium with an harzburgite residue (olivine + orthopyroxene). Also shown is the low-pressure fractional crystallization trend followed under anhydrous conditions from Juster et al. (1989). The contrasting path for fractional crystallization in H<sub>2</sub>O-bearing magma suites (hydrous fractional crystallization) is shown for three different starting compositions: basaltic, primitive basaltic andesite, and a primitive magnesian andesite. Modified after Grove et al. (2003).

The majority of arc magmas erupted on Earth today (i.e., Ewart's average arc magmas) (Figure 9) follow differentiation trends on the  $FeO^*/MgO$  versus SiO<sub>2</sub> diagram that straddle the dividing line between calc-alkaline and tholeiitic, including the Mariana arc, one of Earth's longest arcs. Also shown in Figure 9a are the differentiation trends delineated by the two primitive lavas from the Mt. Shasta, California, region (Table 1). These primary arc magmas lie in a portion of the FeO\*/MgO versus SiO<sub>2</sub> diagram that is not as commonly represented by modern arc environments. The compositions of the lavas from the Mt. Shasta region along with other arc centers, such as the Aleutians in Alaska (Kay 1980, Yogodzinski et al. 1995), the Setouchi volcanic belt in the SW Japan arc (Tatsumi & Ishizaka 1983, Shimoda et al. 1998), and the Chilean arc (Stern & Kilian, 1996) all display higher SiO2 at lower FeO\*/MgO. These lavas represent low temperature, high-H<sub>2</sub>O content, and high-extent melts of the mantle. In addition, their differentiation trends parallel those of the hydrous basaltic melts, but at lower values of FeO\*/MgO. Despite the fact that these compositions are not commonly found in modern arc environments, they represent an important magma type, as many Archean TTG (tonalite-trondhjemite-granodiorite) suites plot in the same portion of the Miyashiro (1974) diagram. Mechanisms that lead to SiO<sub>2</sub> enrichment in these near-primary magmas are not yet understood and represent an area of active research (Wood & Turner 2009).

#### HOW MUCH WATER DO SUBDUCTION ZONE MAGMAS CONTAIN?

Determining the pre-eruptive  $H_2O$  content of a silicate melt from its solidified product remains a challenging task. The solubility of  $H_2O$  is strongly pressure dependent and drops to near zero at atmospheric pressure. Thus, hydrous melts tend to degas as they ascend to the surface. There are several methods for estimating pre-eruptive magmatic H<sub>2</sub>O contents (for a review, see Johnson et al. 1994): indirect measurements using mineral/melt equilibria (Sisson & Grove 1993b), direct measurements of liquids trapped in melt inclusions in minerals (Anderson 1974, Sisson & Layne 1993, Wallace 2005), and estimates of water contents from phase equilibria and/or phenocryst assemblage stability (Moore & Carmichael 1998, Carmichael 2002, Rutherford et al. 1985, Krawczynski et al. 2012). Combined evidence from melt inclusions, phase equilibria, and phenocryst assemblage stability has also been used to track the pre-eruptive history of magma recharge, ascent, and residence in long-lived volcanic systems (Blundy et al. 2006, Schmidt & Grunder, 2011, Johnson et al. 2010). Most of the H<sub>2</sub>O estimates listed in Table 1 were obtained using the phase equilibrium method or from measurements of H<sub>2</sub>O in melt inclusions. Each of these methods is independent, and together, they record a range of  $H_2O$  contents in arc magmas from 0 to 6-7 wt% H<sub>2</sub>O (Figure 10). The value of 6-7 wt% corresponds to vapor saturation of a basaltic magma at relatively shallow depths (6-8 km). This value could indicate that magmas are degassed and originally contained more H<sub>2</sub>O but stalled in a shallow magma chamber.

Thus, still unanswered is the question of the maximum water content present in a subduction zone magma when it last equilibrates with the mantle. The maximum amount of  $H_2O$  that can be incorporated during hydrous flux melting depends on the melt production rate in the mantle wedge and on the temperature structure of the wedge. Above, we proposed that the first melts of the hydrous mantle, derived at the cool base of the wedge, are  $H_2O$  saturated and subsequently ascend into the hotter, shallower overlying wedge where reactive porous flow dilutes the  $H_2O$  in the melt and increases the melting extent. Carmichael (2002) proposed that primary mantle– derived andesite magmas in the Mexican volcanic belt could contain 10 wt%  $H_2O$ . Evolved andesites and dacite lavas at Mt. Shasta contain melt inclusions with >10 wt%  $H_2O$  (Grove et al. 2003), and McCanta et al. (2007) infer 8 wt%  $H_2O$  in the Black Butte dacite. Krawczynski (2011) collected quenched magmatic inclusions from the Mt. Shasta stratocone and from satellite





Temperature and  $H_2O$  contents of the near-primary arc magmas in their mantle source region from **Table 1** (*dark blue squares*) and pre-eruptive temperature and  $H_2O$  contents from mineral melt equilibria for crystallization in the continental crust and uppermost mantle. Porphyritic (*open circles*) and aphanitic lavas (*filled circles*) from the south Cascades region (Sisson & Grove 1993b) are shown in the gray field and highlight the range of  $H_2O$  contents in basaltic melts that crystallize at shallow crustal pressures. The dark yellow diamonds (Krawczynski et al. 2012) represent  $H_2O$ -saturated melts of andesite that crystallized in a magmatic crystal mush column below Mt. Shasta, California, that extends to mantle depths. A challenge is to understand the conditions that lead to the development of the high- $H_2O$ -low-temperature melts near the mantle-crust transition below some arc volcanoes. Abbreviations: BA, basaltic andesite, HAB, high-alumina olivine tholeiite; wt%, weight percent.

vents and found that primitive mantle melts were emplaced into the Mt. Shasta plumbing system with  $\sim$ 14 wt% H<sub>2</sub>O ( $\sim$ 800 MPa H<sub>2</sub>O saturated) and subsequently crystallized and degassed as they ascended through the crust (**Figure 10**). The Mt. Shasta inclusions contain the assemblage of high Mg# pargasitic amphibole + olivine + pyroxene that requires high H<sub>2</sub>O pressures to stabilize it on the liquidus; Krawczynski et al. (2012) reproduced this liquidus assemblage in H<sub>2</sub>O-saturated experiments on Mt. Shasta andesite liquids that are in equilibrium with mantle peridotite. Ridolfi et al. (2010) found similarly high pressures and H<sub>2</sub>O contents preserved in the phenocryst assemblages in andesite and dacite lavas from Mt. St. Helens, Washington; Soufriere Hills, West Indies; Redoubt, Alaska; and El Reventador, Ecuador.

Also plotted in **Figure 10** are the temperature and  $H_2O$  contents inferred for the primitive arc magmas from **Table 1**. These liquids plot at higher temperatures and span the water contents estimated from more evolved arc basalts. The trend of the primitive arc magmas extends up to 7.5 wt%. If we extend the trend to the maximum water contents calculated for the Mt. Shasta inclusions, several of which are primitive magmas, we infer that primitive arc magmas could contain

#### **Primitive magma:**

the most compatible element-rich compositions in a given suite; refers to near-primary magmas up to 14 wt%  $H_2O$  when separated at the top of the mantle wedge. Mantle melts with higher  $H_2O$  contents could form by reactive crystallization at the upper colder portion of the mantle wedge where temperature decreases with decreasing depth. Alternatively, water input into the mantle wedge may be a variable process, and these relatively cold but very high  $H_2O$  content rocks reflect ascent paths through a colder part of the wedge (i.e., on the forearc side of path B, **Figure 6**) and thus experienced variable wedge flux melting. The positive sloping trend in  $H_2O$  content and temperature shown for the Mt. Shasta inclusions in **Figure 10** likely records continual vapor saturation of the highest water content magmas as they ascend through a crystal mush column prior to eruption from the Mt. Shasta stratocone (Krawczynski 2011).

#### THE LOCATION OF ARC VOLCANOES

The classic view of the localization of volcanic arcs is that they occur where the depth to the slabmantle wedge boundary is  $110 \pm 20$  km (Jarrard 1986). Several recent studies provide excellent global compilations regarding the locations of arc volcanoes and the corresponding plate tectonic observables (i.e., convergence rate, slab dip, slab age, etc.) that can be examined for systematic variations (England et al. 2004, Syracuse & Abers 2006). These global compilations demonstrate that the positions of volcanoes in arcs vary in relation to the plate tectonic observables. England & Katz (2010a, 2010b) used these observations to suggest a model where arc volcanoes are located above the place in the mantle wedge where the anhydrous peridotite solidus makes its closest approach to the trench. However, the petrologic constraints on hydrous magma generation at volcanic arcs discussed in this paper, including the H<sub>2</sub>O content of arc magmas, the effects of  $H_2O$  on basaltic-andesitic phase equilibria, the P-T conditions of hydrous mineral breakdown in the subducting lithosphere, and the P-T conditions of the H<sub>2</sub>O-saturated solidus, all suggest H<sub>2</sub>O plays a critical role in the generation of primitive arc magmas. Grove et al. (2009) proposed that the locations of arc volcanoes in these global compilations are the result of the interplay between the location of H<sub>2</sub>O-rich fluid release from hydrous minerals in the subducted oceanic lithosphere and the temperature of the mantle wedge above this region of fluid release.

Specifically, the interactions of the petrologic variables (e.g., the P-T slope of the vaporsaturated peridotite solidus and hydrous mineral breakdown) with the plate tectonic variables (i.e., slab dip and convergence rate) appear to control the locations of volcanoes above subduction zones, rather than a uniform depth to the slab (Figure 11). As shown in Figure 6, melting in the mantle wedge can initiate just above the slab over a range of depths (illustrated between paths A and C, Figure 6a). Recall that not all these initial melts will evolve in a way that leads to extensive melting and eruption from an arc volcano on Earth's surface. Melts produced from H<sub>2</sub>Osaturated melting at path A will likely freeze as soon as they ascend into the cooler overlying mantle (Figure 6a). Where the overlying mantle is higher in temperature (e.g., path B, Figure 6a), the initial H<sub>2</sub>O-rich melts formed at the slab-wedge interface can ascend and continue to melt the mantle by flux melting at  $H_2O$ -undersaturated conditions (Grove et al. 2006). These melts then become hot enough and of sufficient volume to ascend through the shallower, cooler part of the mantle and crust and erupt at arc volcanoes. Path C marks the termination of H<sub>2</sub>O-rich fluid supply from the breakdown of chlorite and/or serpentine, which will cut off H<sub>2</sub>O-saturated melting in the mantle wedge. Between paths B and C, the overlying mantle will be hot enough to allow flux melting to produce hot H<sub>2</sub>O-undersaturated melts in the shallower hot core of the mantle wedge that can feed the growth of arc volcanoes on Earth's surface.

Grove et al. (2009) undertook numerical simulations where slab dip (32–64°) and convergence rate (20–100 mm/year), which are observed to vary in subduction zones globally, are used to estimate the depth and temperature at the slab-wedge interface. These models were used to assess



The relationship between the plate tectonic variables at global arcs including (*a*) dip of slab plotted against arc-trench distance (modified from Grove et al. 2010) and (*b*) dip of slab plotted against depth to slab beneath the volcanic front and the corresponding predictions from the numerical models of Grove et al. (2009). The location of volcanoes globally in arc systems can be explained by the interplay of the petrologic variables (e.g., temperature of the H<sub>2</sub>O-saturated solidus, pressure-temperature conditions of hydrous mineral breakdown) and the plate tectonic variables at subduction zones (e.g., slab dip, convergence rate), as exemplified by Paths A–C from **Figure 6**. Dashed line in panel *b* illustrates the role slab dip may play in locating the shallowest depth of melting beneath arc volcanoes. Volcanoes that form by hydrous flux melting should plot between paths A and C in **Figure 6**.

the best fit geometry of the A to C region at global subduction zones where the P-T conditions are above the H<sub>2</sub>O-saturated solidus and within chlorite stability, thus providing the supply of H<sub>2</sub>O to trigger melting at the H<sub>2</sub>O-saturated solidus (P,T<sub>melt</sub> in **Figure 6***b*). The models assume that the shallowest depth of melting, as intersected by path A, is at the low-pressure intersection of chlorite stability and the H<sub>2</sub>O-saturated solidus (lowest pressure point in the region of P,T<sub>melt</sub>). Path B is designated as the vertical path that intersects the nose of the 1250°C isotherm in the models (**Figure 11**). Path C is located at the high-pressure intersection of chlorite stability and the H<sub>2</sub>O-saturated solidus (highest pressure point in the region of P,T<sub>melt</sub>). When the distance from the trench to the projected intersection of paths A, B, and C with the surface is compared with the observed arc-trench distances from the global subduction data of England et al. (2004), the models match the observed variations reasonably well (**Figure 11***a*). Thus the model results illustrate that arc volcanoes form between paths B and C as predicted (where both H<sub>2</sub>O supply and temperatures are high). The arc-trench distances in the global compilation of Syracuse & Abers (2006) tend to be 50–100 km greater than those determined by England et al. (2004), which likely stems from systematic differences in how these two studies located the exact position of the trench. In **Figure 11***b*, the depth to the slab at paths A, B, and C from the numerical models are compared with the depth to the slab below arc volcanoes determined in the global compilations. The depth to slab below arc volcanoes varies between ~65–70 km and more than 160 km. Slab dip does not appear to be a strong control on the depth to the slab below arc volcanoes, although it may play a greater role in determining the shallowest depth of melting (illustrated by the negative slope of the dashed line in **Figure 11***b*). Instead, convergence rate likely plays a role in the positions



#### Figure 12

Pressure-temperature (P-T) diagram showing the temperature at the slab-wedge interface predicted by numerical models of Syracuse et al. (2010) for the assumption of full coupling between the slab and the mantle wedge at 80-km depth. Illustrated are the range of slab surface P-T conditions predicted by the models of Syracuse et al. (2010) (*red lines*) and the P-T conditions determined for the slab-wedge interface below modern arcs (*gray box*). The H<sub>2</sub>O-saturated peridotite solidus (*blue curve*) and hydrous phase stability fields (*dark gray curves*) from **Figure 6** are superimposed on the models along with the experimentally determined melting curve for H<sub>2</sub>O-saturated ocean crust (Liu et al. 1996) (*purple curve*) and sediment (Nichols et al. 1994) (*dark yellow curve*). Also superimposed on this plot are the phase relations for hydrous fertile peridotite from **Figure 4** showing chlorite dehydration, the H<sub>2</sub>O-saturated melting curves and hydrous-phase breakdown reactions for H<sub>2</sub>O-rich mantle minerals overlap with the thermal model estimates of the conditions at the base of the mantle wedge and interface with the subducted oceanic lithosphere. We suggest that this leads to the initiation of hydrous flux melting.

of arc volcanoes and paths A, B, and C, as the distance of any given isotherm from the trench decreases with increasing convergence rate. Thus, convergence rate appears to play a greater role in the location of arc volcanoes, and slab dip plays a greater role in determining the location where melting begins in the mantle wedge closest to the trench.

The recent geodynamic models of Syracuse et al. (2010) also predict the thermal structure of the mantle wedge–subducted lithosphere interface. Their models show how the temperature structure for the global range of plate tectonic variables observed in arcs is influenced for different assumptions of the nature of the coupling between the slab and mantle wedge. **Figure 12** shows the results of one of their models in which full coupling between the slab and wedge occurs at 80-km depth. The plate tectonic variables that characterize each arc in their data set are used to calculate a temperature-depth profile for the surface of the subducting slab at that arc. The coincidence of the slab surface temperature with the vapor-saturated solidus and chlorite breakdown for the majority of the models supports the notions that  $H_2O$ -saturated peridotite melting can be triggered near the slab-wedge interface and that it happens at a depth in the mantle that coincides with the depth to the slab-wedge interface observed below arcs.

#### SUMMARY POINTS

- 1. Near-primary magmas are erupted in many arcs worldwide and require  $H_2O$  contents that range from 0 to 7.5 wt% to be in equilibrium with a lherzolite or harzburgite assemblage at their liquidi.
- P-T estimates from metamorphic rocks and thermal models agree that the slab-wedge interface is ~700–900°C beneath the volcanic zones at arcs. These temperatures put critical constraints on the stability, or lack thereof, of hydrous minerals at the location of H<sub>2</sub>O release in the mantle wedge.
- 3. Small amounts of H<sub>2</sub>O given off by the subducting slab lower the melting point of peridotite to between 800 and 820°C at 3 GPa. As the initially H<sub>2</sub>O-saturated melt ascends in the wedge, it constantly dissolves silicate material by reactive flux melting in a process that increases the melting extent and dilutes the H<sub>2</sub>O content of the melt.
- 4. The process of reactive flux melting, combined with adiabatic decompression and other possible melting and crystallization processes in the mantle wedge, creates a spectrum of near-primary melts injected into the crustal plumbing systems at arc volcanoes. These melts range from dry olivine tholeiites to hydrous andesites with Mg#s in excess of 0.7.
- 5. Arc magmas have distinct chemical variations as they crystallize owing to the presence of H<sub>2</sub>O. The distinctive enrichment of SiO<sub>2</sub> in these magmas is created by the early crystallization of olivine, pyroxene and amphibole. In addition, the crystallization of plagioclase is suppressed at mid- to lower crustal pressures in H<sub>2</sub>O-rich magmas. When it appears the plagioclase is high-Ca and high-Al. Phase equilibria and melt inclusion studies suggest that magmas with 6–8 wt% H<sub>2</sub>O are common in the crust, and upper limits of pre-eruptive H<sub>2</sub>O contents of magmas are ~12–14 wt%.
- 6. The stability of hydrous minerals in the mantle wedge and the ability for H<sub>2</sub>O-rich melts to travel through the mantle wedge are affected by the plate tectonic variables of subduction zones. These constraints together with the temperature structure and mantle-flow patterns beneath arcs can be used to describe the location of arc volcanoes.

#### **FUTURE ISSUES**

- 1. The fate of mantle melts as they enter the cooler, shallower mantle wedge above the hot core of the subduction zone remains uncertain. At present, we lack a clear understanding of the relative roles of advection of heat and melt into this region and their consequences.
- The nature of the Moho and lithosphere-asthenosphere boundary beneath arcs raises similar questions. Improvements in numerical models, observations of the seismic structure below arcs, and synthesis of the evidence from different disciplines are needed.
- 3. The observed diversity in near-primary mantle melts in arcs has yet to be adequately explained. Better sampling of the compositional diversity and its spatial distribution in the global arc system and the relation of this compositional variability to the plate tectonic variables may lead to new insights into the evolution of arc magmatic products through geologic history.
- 4. The nature of material transport from the subducted oceanic lithosphere into the overlying mantle wedge is very poorly constrained. We lack an understanding on the physical controls on this process. Geologic evidence from exhumed arc roots can provide key constraints. Also lacking is an understanding of the most important material transport and chemical reactions that occur between slab and wedge. For example, the influence of variations in H<sub>2</sub>O and its spatial distribution in the slab and wedge are largely unknown, as is the relationship of fluid additions to the abundances of fluid-mobile elements in arc magmas.

#### **DISCLOSURE STATEMENT**

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